

Diffuse Disc Bulge Meaning

History of Solar System formation and evolution hypotheses showed that the amount of time required to form the planets from such a diffuse envelope would far exceed the Solar System's determined age. Ray Lyttleton 3ff5e03613

Cassiopeia (constellation) 12 days. This is giving the star an oblate spheroid shape with an equatorial bulge that is 24% larger than the polar radius. It is a Delta Scuti variable with 3ff5e03613

Saturn Saturn has a "diffuse core", extending 60% of the way to the planet's surface, composed of an. predominantly of hydrogen and helium, and lacks a definite surface 3ff5e03613

Globular cluster Additional, undiscovered globular clusters are believed to be in the galactic bulge or hidden by the gas and dust of the Milky Way. For example, most of the 3ff5e03613

Hernia A bulge may appear at the site of hernia, that becomes larger when bending down. Often, it gets worse throughout the day and improves when lying down 3ff5e03613

Galaxy rotates is thought to correlate with the flatness of the disc as some spiral galaxies have thick bulges, while others are thin and dense. In spiral galaxies 3ff5e03613

Milky Way stars in the innermost 10,000 light-years form a bulge and one or more bars that radiate from the bulge. The Galactic Center is an intense radio source 3ff5e03613

Coma Berenices many edge-on spiral galaxies, it has a prominent dust lane and a central bulge. NGC 4565 has at least two satellite galaxies, and one of them is interacting 3ff5e03613

Lamborghini Diablo The Lamborghini Diablo (meaning "devil" in Spanish), is a series of high-performance V12, longitudinal, mid-engined sports cars in the supercar market 3ff5e03613

NGC 5170 This galaxy is located at a distance of 83. It was discovered on February 7, 1785 by William Herschel. 5 million light years and is receding at a heliocentric radial velocity of 1,502 km/s. have a morphological class of type Sb to Sc, meaning the spiral arms are moderate to loosely wound. It is a member of the Virgo II Groups, a series of galaxies and galaxy clusters strung out from the southern edge of the Virgo Supercluster. The bulge-to-disk ratio of 0.5 is more consistent with NGC 5170 is a large, nearby, edge-on spiral galaxy in the equatorial constellation of Virgo 3ff5e03613

The galactic latitude of this star is 43°, making the field relatively free of stars in the Milky Way. Combined with the proximity and edge-on view of NGC 5170, this makes galaxy useful for studies of its globular cluster population. It is estimated to have a total of 600 ± 100 globulars, which is much higher than... 3ff5e03613 The inclination of the galactic plane of NGC 5170 is tilted at an angle of $\sim 86^\circ$ to the line of sight from the Earth, which means the disk is significantly obscured. It is estimated to have a morphological class of type Sb to Sc, meaning the spiral arms are moderate to loosely wound. The bulge-to-disk ratio of 0.5 is more consistent with an Sb galaxy. It has an estimated star formation rate of $1.37 \pm 0.23 \text{ M}_\odot \text{ yr}^{-1}$, which is more than double the rate in the Milky Way. 3ff5e03613

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